

ASTR 469

HW #6

a) $\lambda \approx 80^\circ W$ $t = 15.5$
 $LST = 3.45^h = 3^h 27^m$

b) $+1^h$

c) $\sin(\alpha) = \sin(\delta) \sin(\varphi) + \cos(\delta) \cos(HA) \cos(\varphi)$
 $\varphi \hat{=} +38.5^\circ$ $\delta \hat{=} +63^\circ$
 $\Rightarrow \alpha \hat{=} 65^\circ$

d) $\sim 60^\circ$

e) $1 \text{ GHz} \approx 30 \text{ cm}$ $2 \text{ GHz} \approx 15 \text{ cm}$

$$\frac{1}{\lambda} = R \left(\frac{1}{n^2} - \frac{1}{(n+1)^2} \right)$$

$$R = 1.1 \times 10^7 \text{ m}^{-1}$$

Solves to $n = 187 - 148$

2a) W3Main: Peak: 0.8K, FWHM: ~ 20 km/s
Velocity: ~ -40 km/s
W3OH: Peak: 0.15K, FWHM: ~ 30 km/s
Velocity: ~ -45 km/s

b) The continuum level is key here.

W3Main has $T_c \sim 80$ K.

W3OH has $T_c \sim 12$ K.

H I₁ and H I₂ have $T_c \sim 0$ K

For the two W3 positions, H I foreground to W3 is seen in absorption. If T_c is higher, absorption is stronger. For two H I positions, no background continuum, so no absorption.

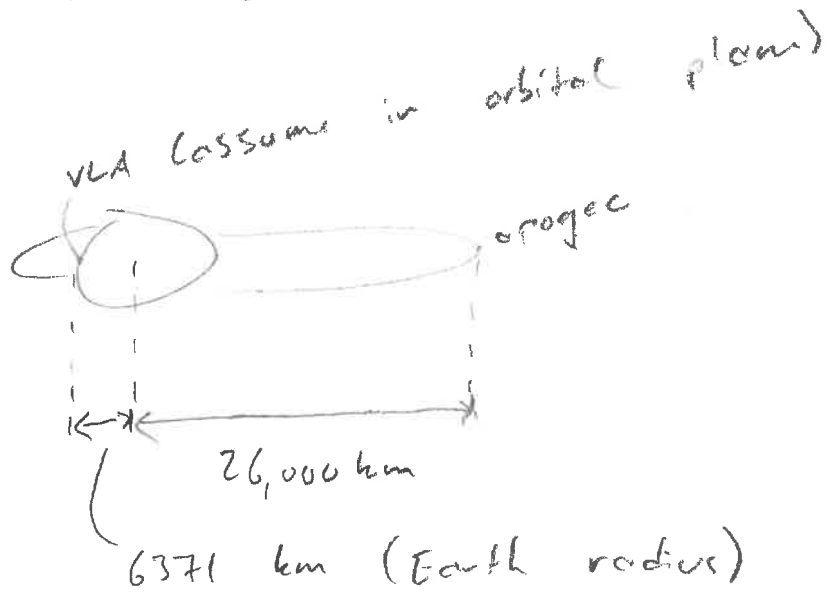
$$3) \lambda_{\max} = \frac{0.002898 \text{ m}}{T}$$

$$56 \text{ Hz} \hat{=} 0.06 \text{ m}$$

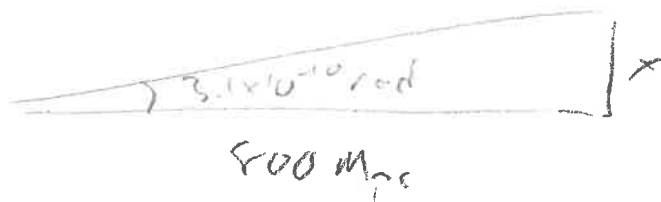
$$\Rightarrow T \hat{=} \frac{0.003}{0.06} = 0.05 \text{ K}$$

Since the universe is at least $\sim 3 \text{ K}$,
nothing peaks in the radio.

41) $\nu = 30 \text{ GHz}$ $\lambda = 0.01 \text{ m}$



$$\theta = \frac{127 \lambda}{D} \approx \frac{0.01}{(26000 + 6400) \times 10^3} = 3.1 \times 10^{-10} \text{ rad}$$



$$\Rightarrow 3.1 \times 10^{-10} = \frac{x}{800 \text{ Mpc}}$$

$$\Rightarrow x = 2.5 \times 10^{-7} \text{ Mpc} = 0.25 \text{ pc}$$