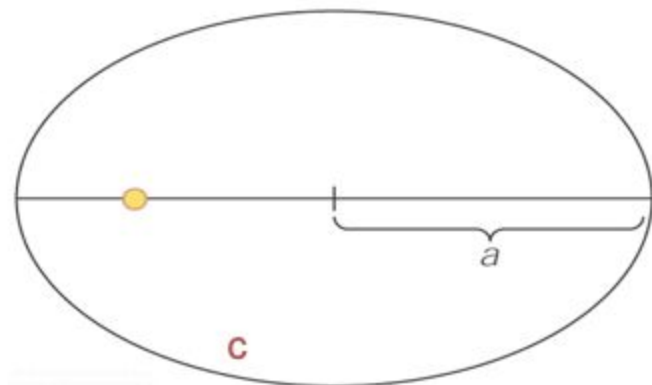


# ASTR367

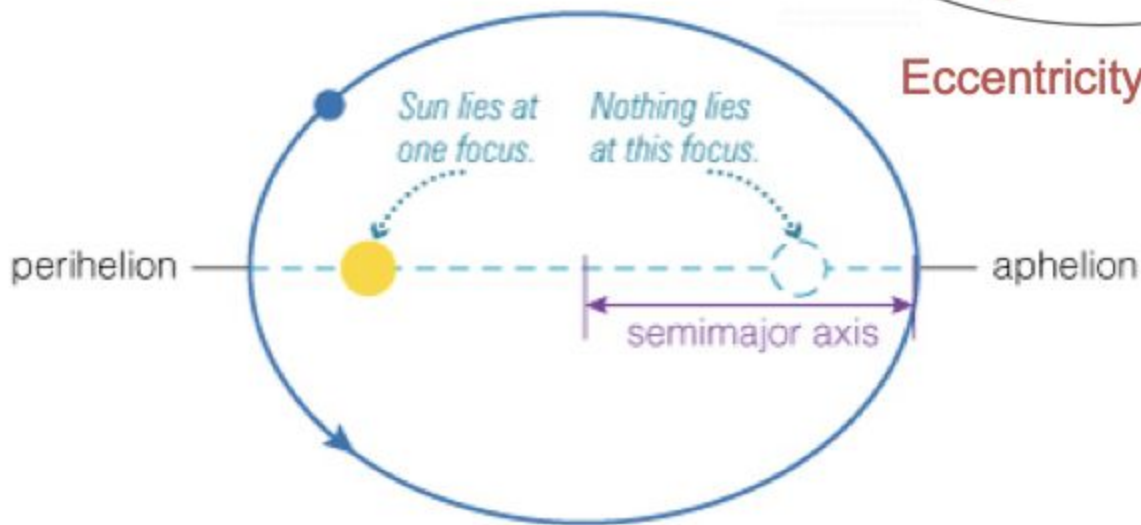
Binaries

# Kepler's 1<sup>st</sup> Law

**Law #1:** The orbits of the planets are **ellipses** with the Sun at one focus.



$$\text{Eccentricity } e = c/a$$



# Eccentricities of Ellipses

1)

$$e = 0.02$$

2)

$$e = 0.1$$

3)

$$e = 0.2$$

4)

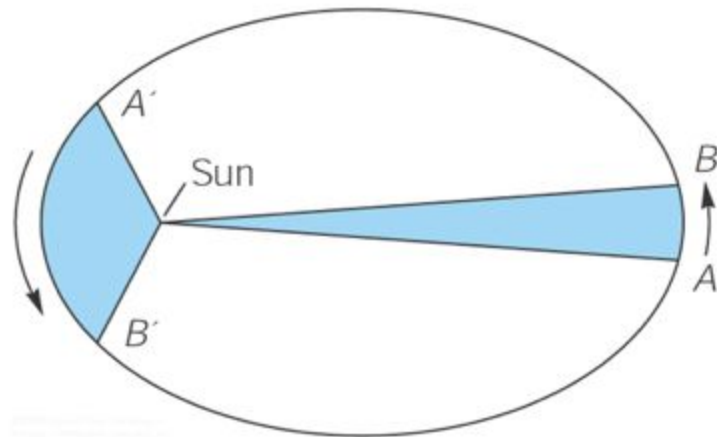
$$e = 0.4$$

5)

$$e = 0.6$$

## Kepler's 2<sup>nd</sup> Law

**Law #2:** A line from a planet to the sun sweeps over equal areas in equal intervals of time.



## Kepler's 3<sup>rd</sup> Law

A planet's orbital period (P) squared is proportional to its average distance from the sun (a) cubed:

**Closer orbits go faster**

$$P_y^2 = a_{AU}^3$$

( $P_y$  = period in years;  
 $a_{AU}$  = distance in AU)

<http://www.solarsystemscope.com/>

<http://astro.unl.edu/naap/pos/animations/kepler.html>

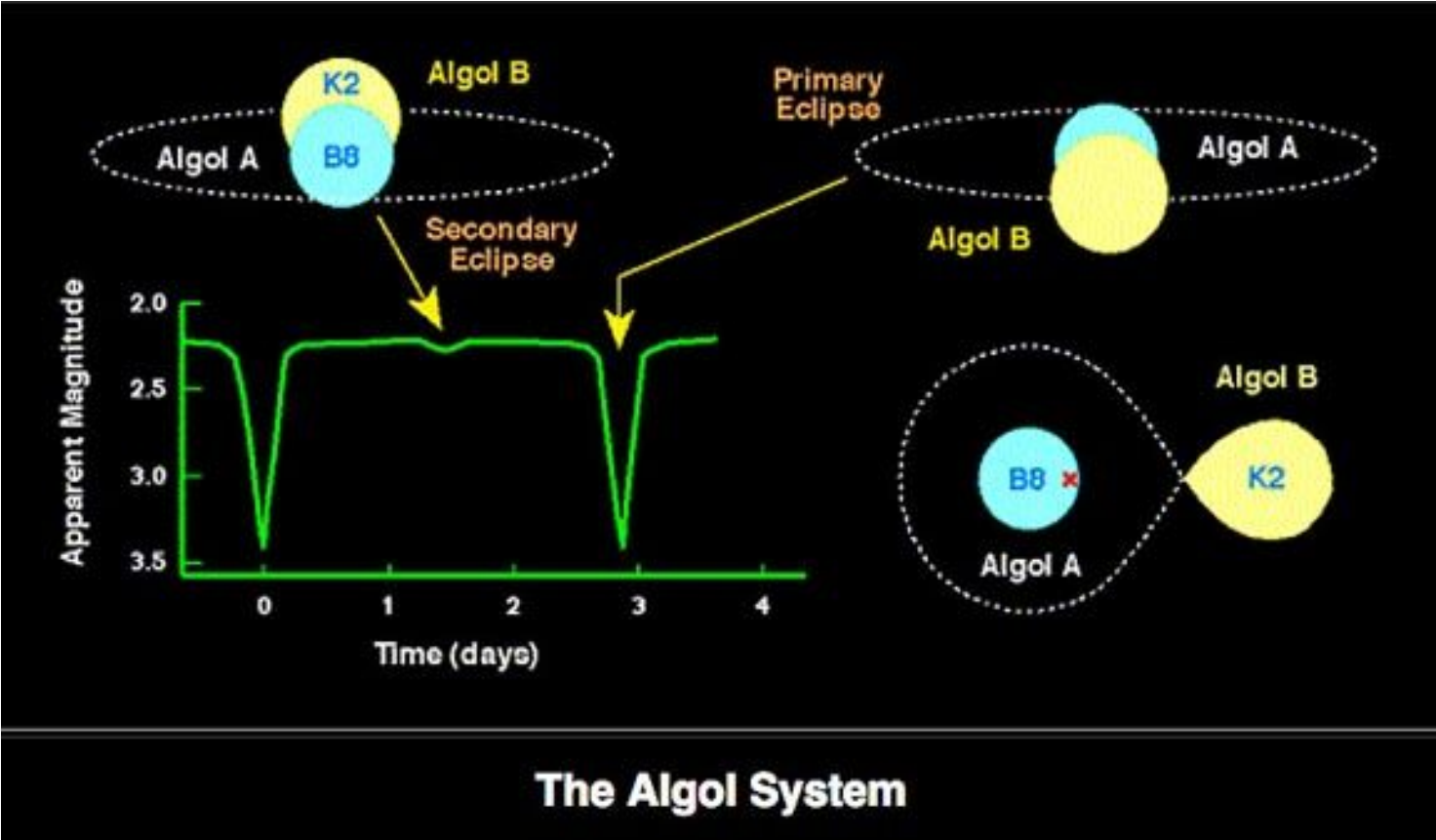
# Optical doubles



# Visual Binaries

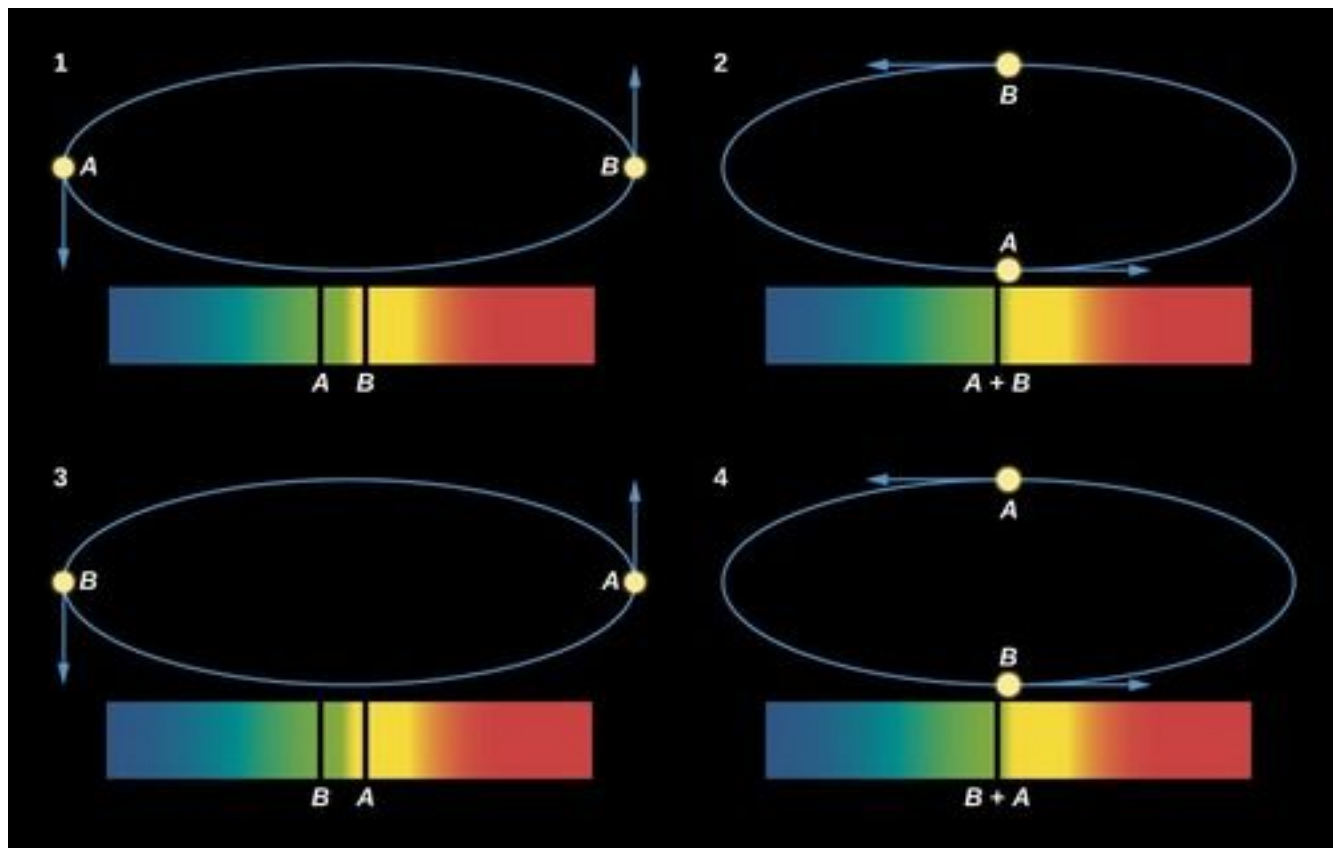


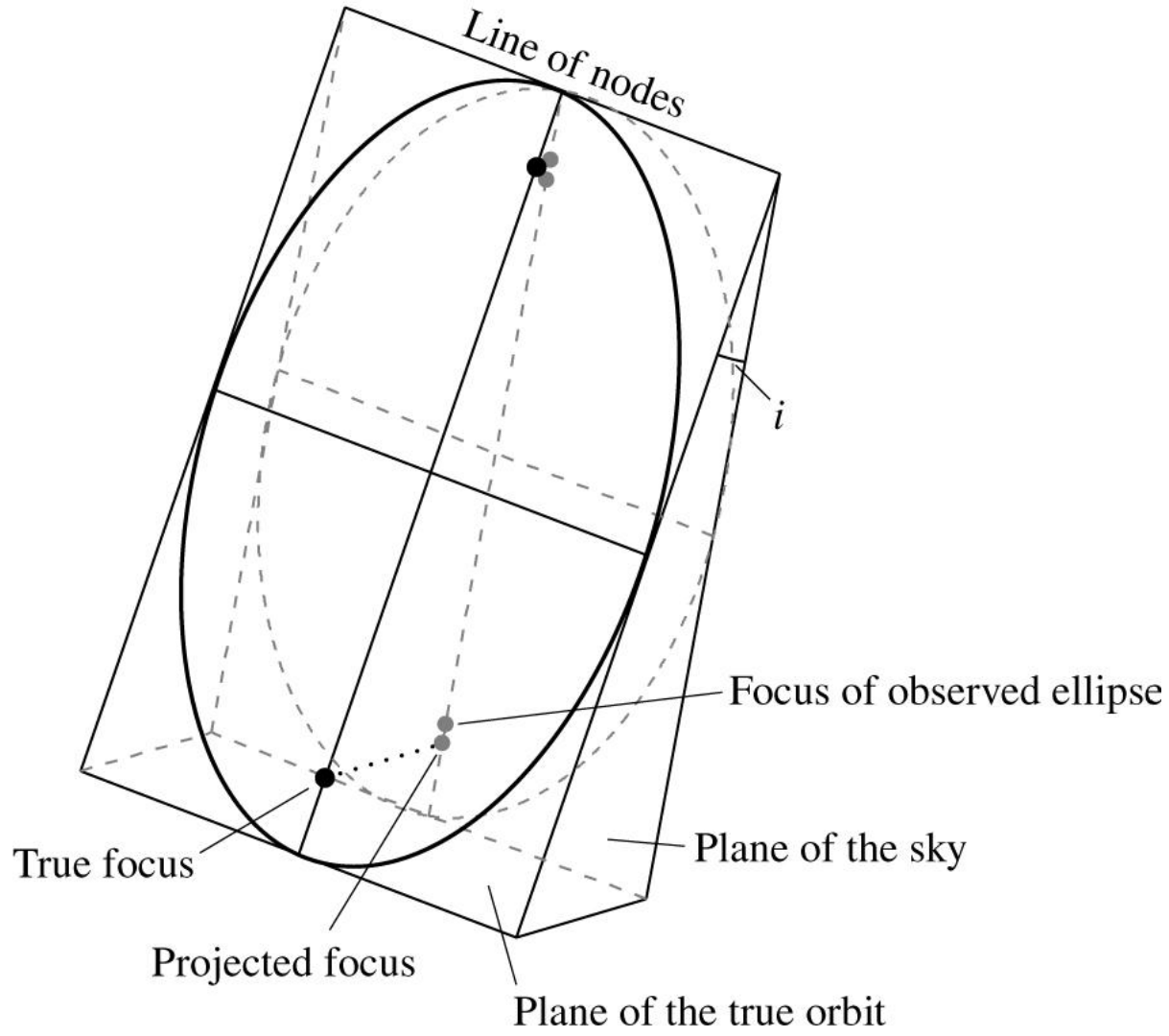
# Eclipsing binary

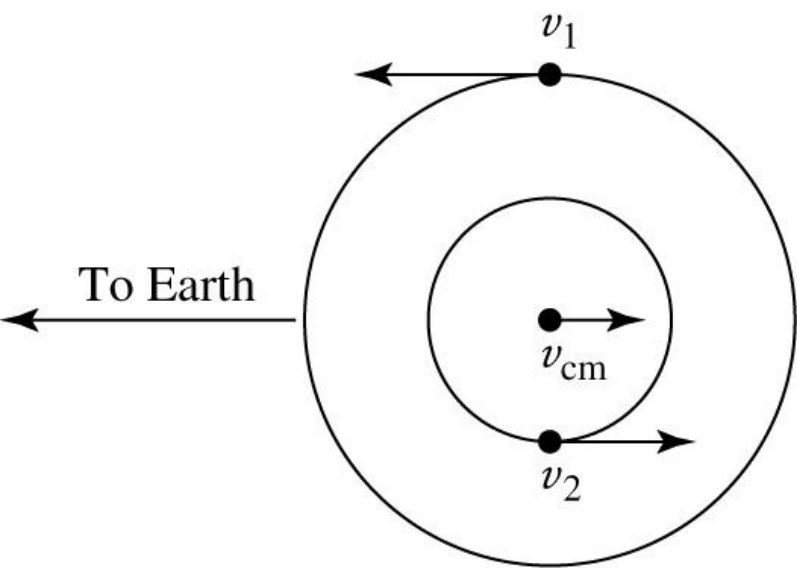




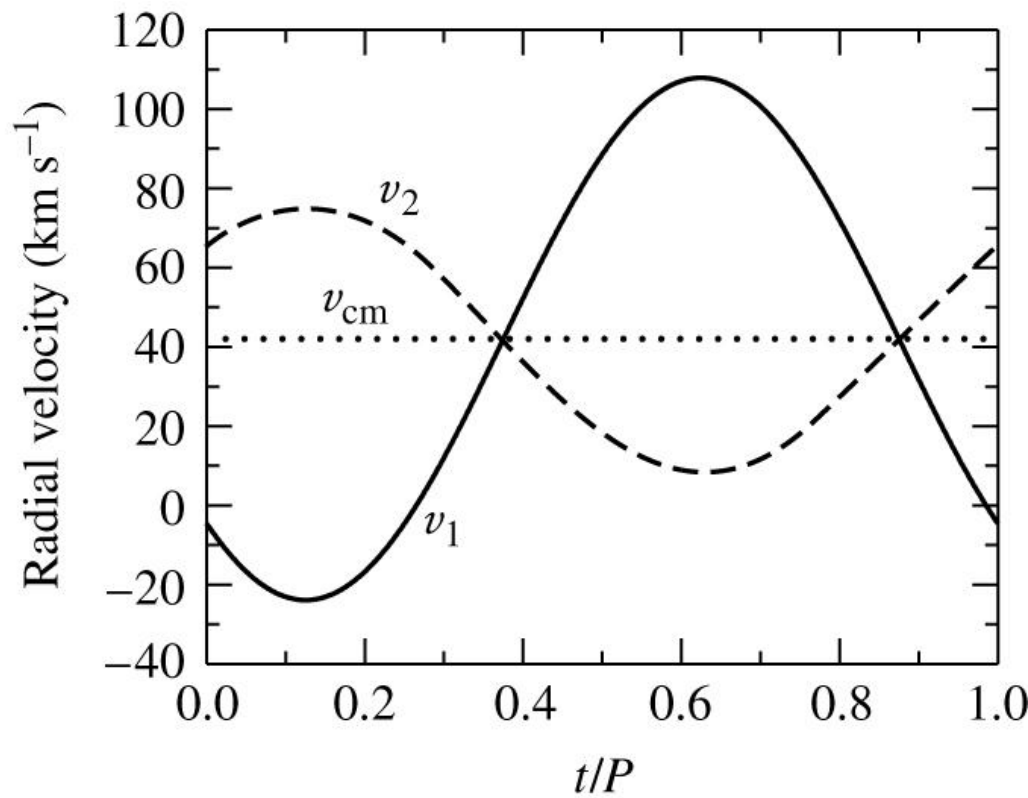
# Spectrum Binaries



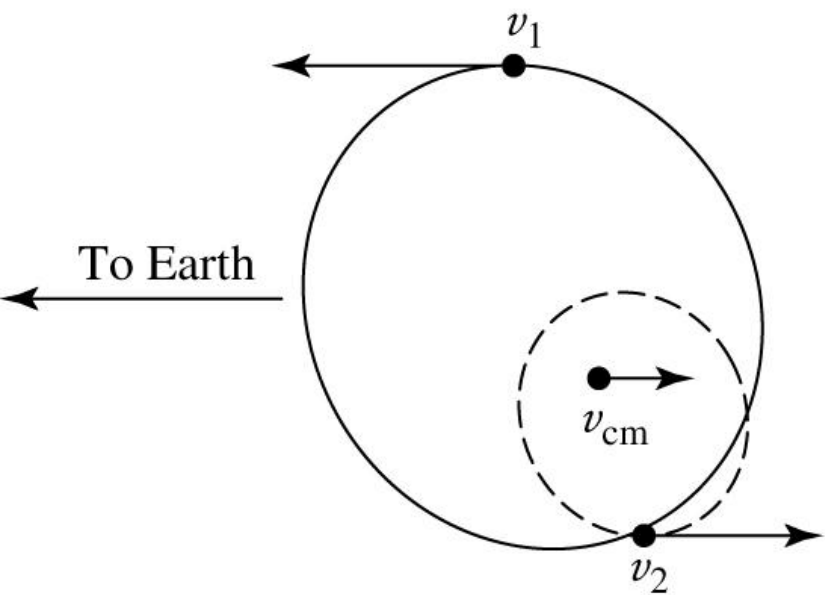




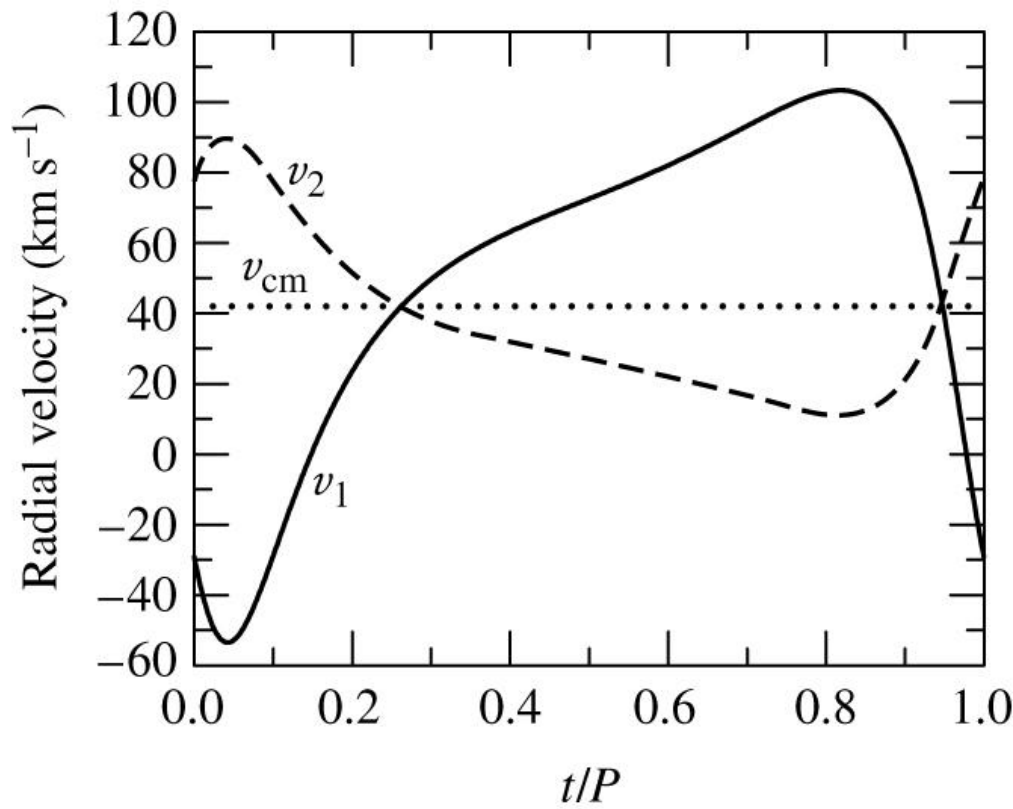
(a)



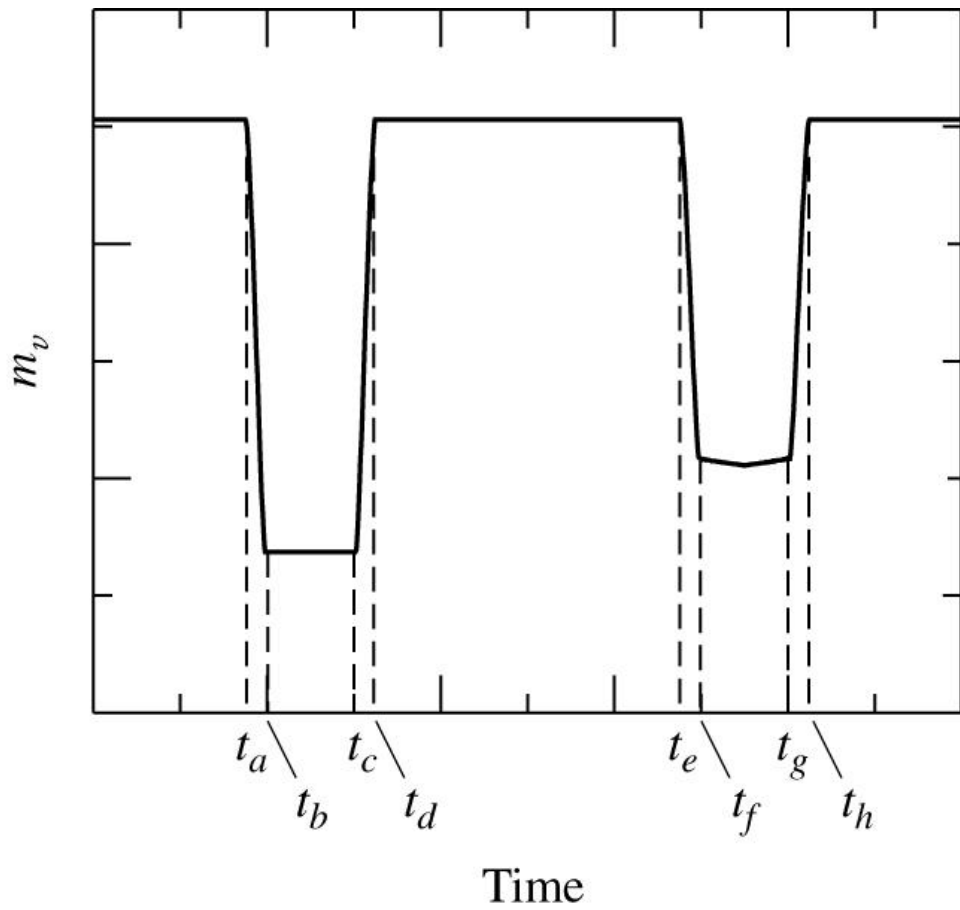
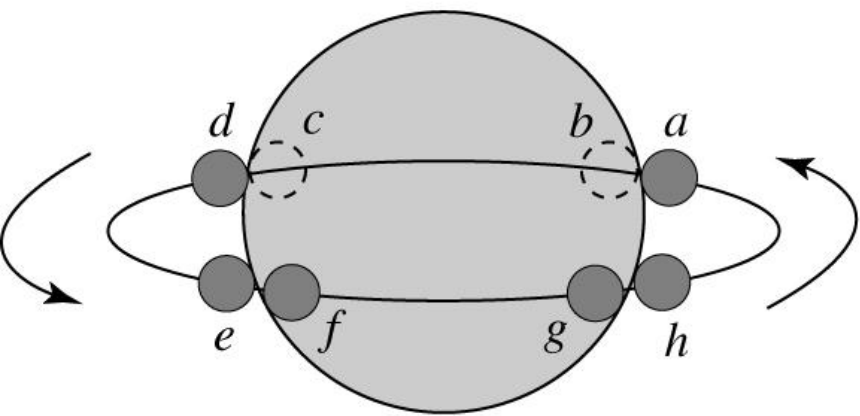
(b)

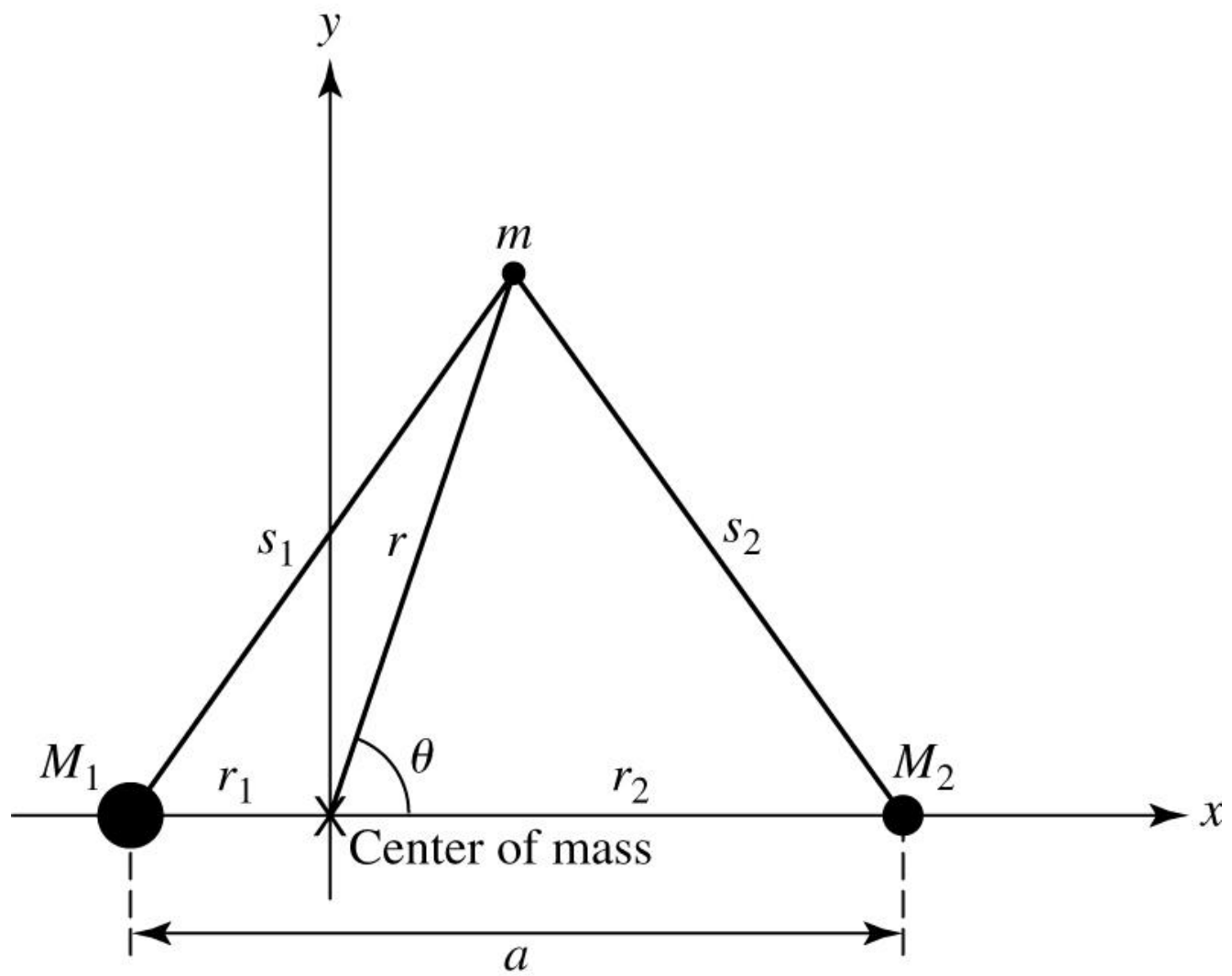


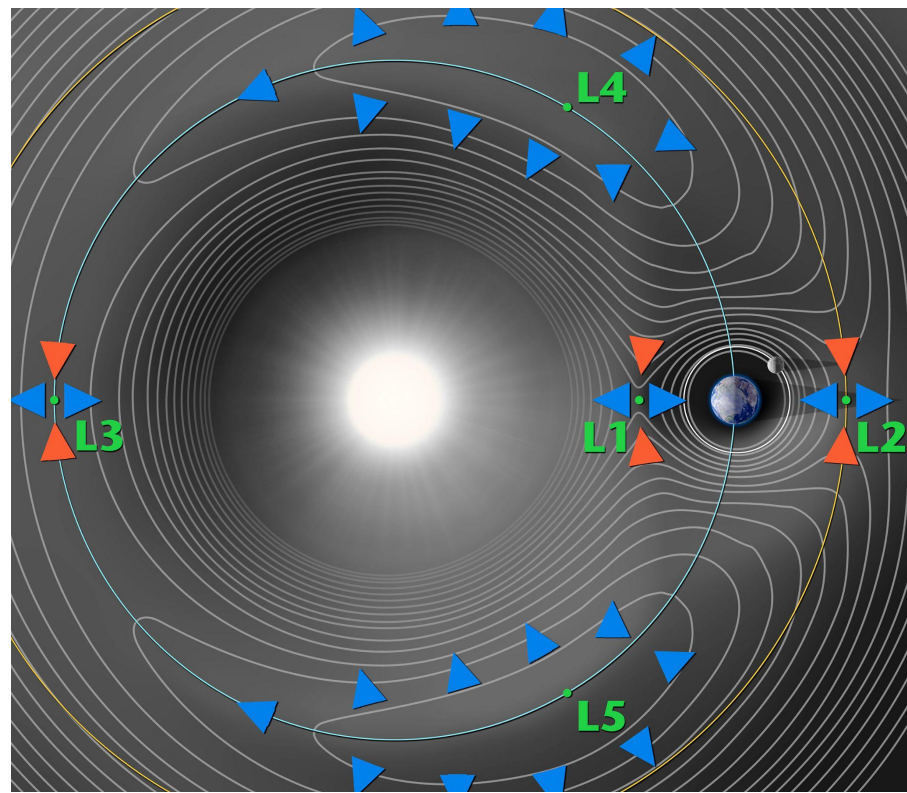
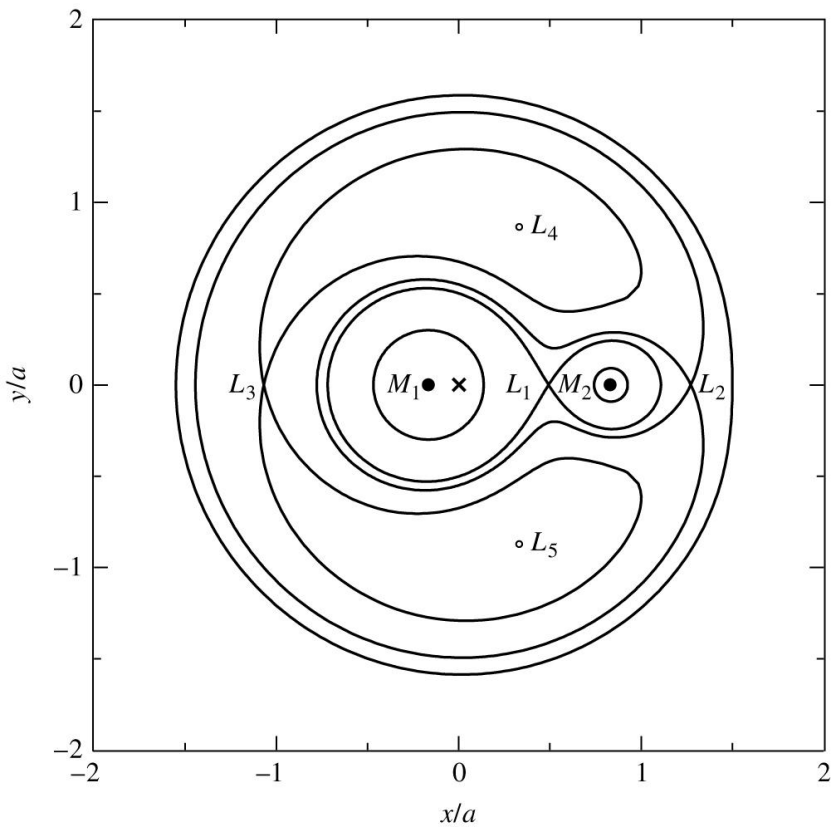
(a)

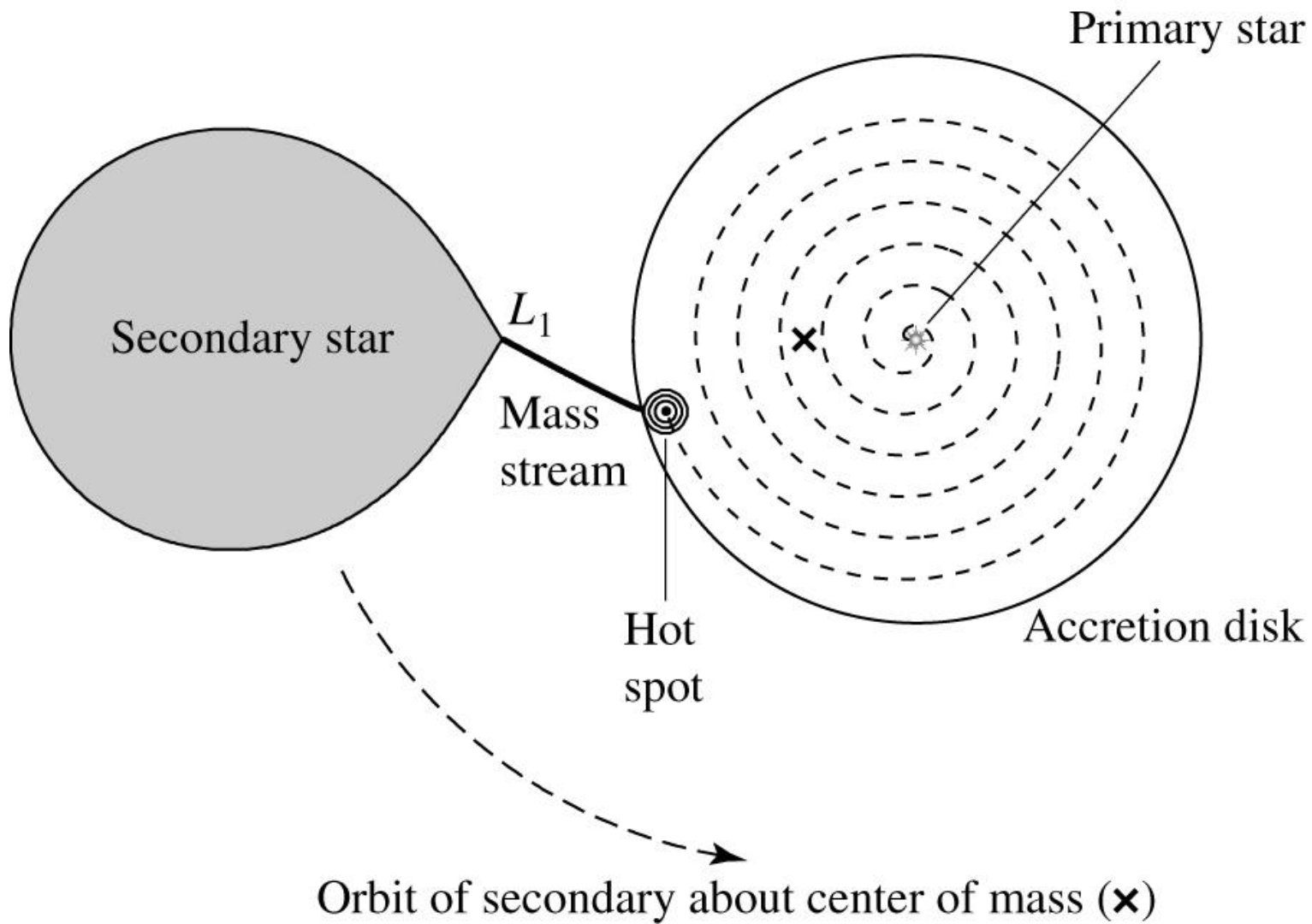


(b)

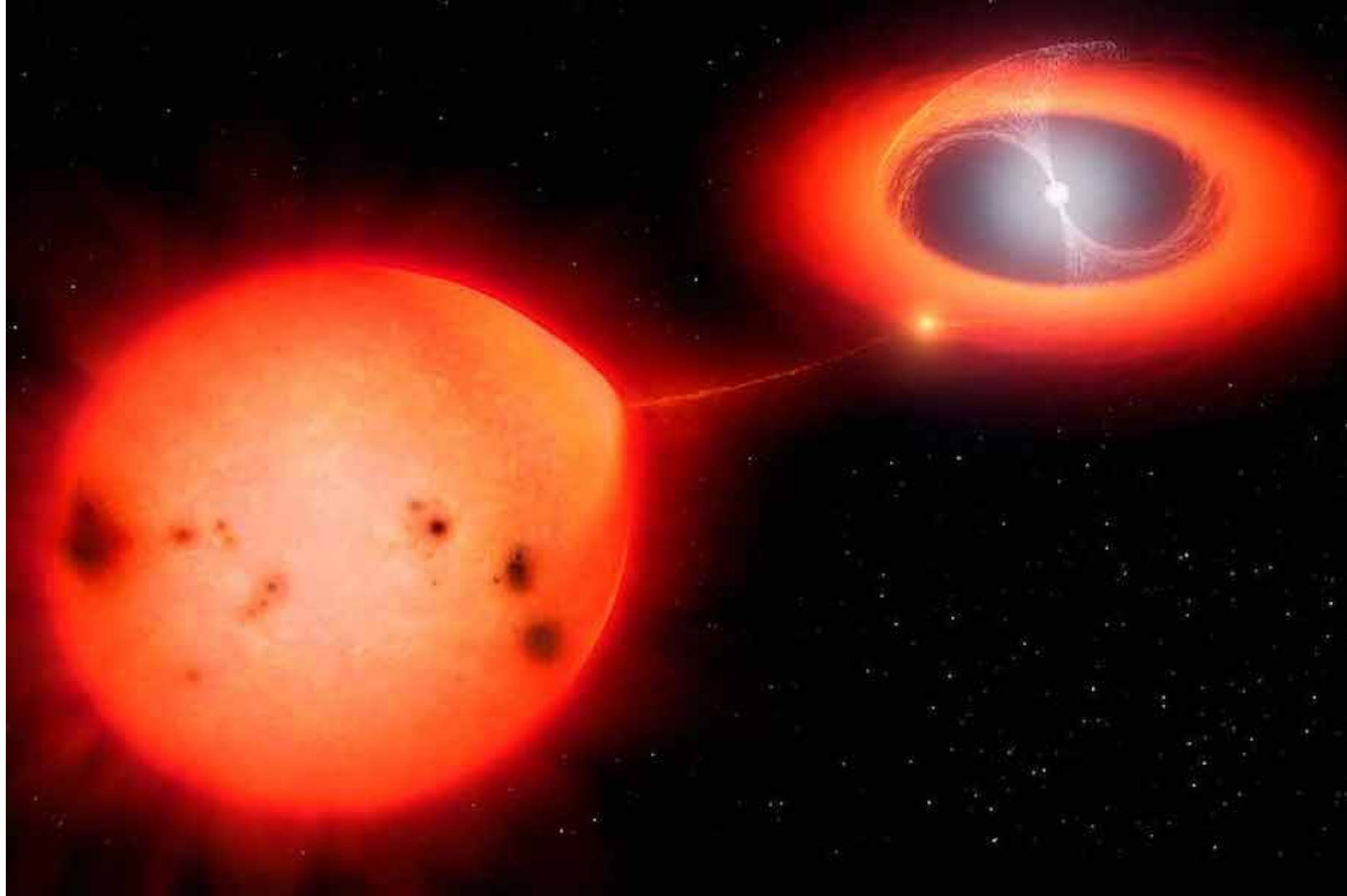


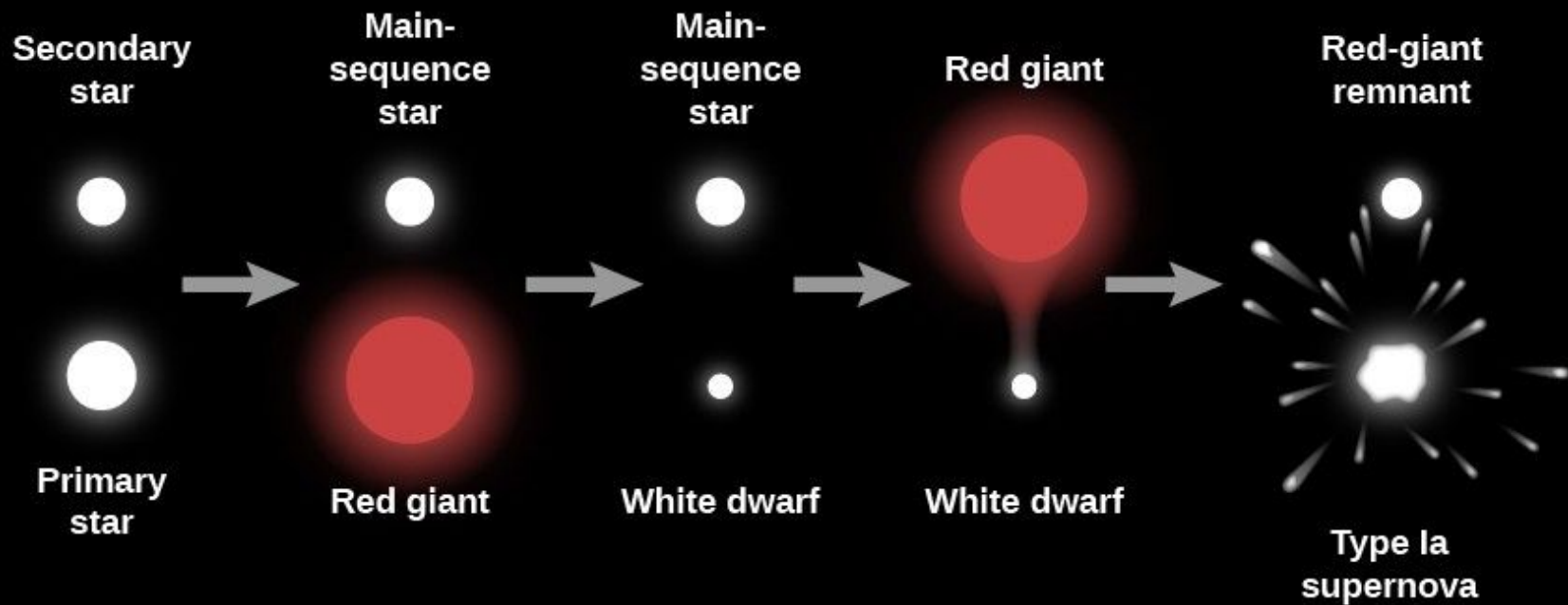












*Not to scale*

