

ASTR 705

HW #2

$$1a) \quad \frac{n_u}{n_l} = \frac{g_u}{g_l} e^{-h\nu_{ul}/kT_s} = 3 e^{-h\nu_{ul}/kT_s}$$

b) Let's look at the exponent

$$\nu_{ul} = 1420 \text{ MHz} \quad (= 2 \text{ cm})$$

$$\Rightarrow \frac{h\nu}{k} = 0.07 \text{ K}$$

$$n(\text{HI}) = n_u + n_l = n_u + \frac{n_u}{3} e^{h\nu_{ul}/kT_s}$$

$$= \frac{4}{3} n_u e^{\frac{0.07}{T_s}}$$

$$\text{If } T_s = 100 \text{ K, } n(\text{HI}) = \frac{4}{3} n_u e^{0.07/100} \\ = 1.0007 \cdot \frac{4}{3} n_u$$

$$c) \text{ If } T_s = 20 \text{ K, } n(\text{HI}) = 1.003 \cdot \frac{4}{3} n_u$$



$$2) \text{ Derive } e^{-h\nu/kT_{ex}} = \frac{e^{-h\nu/kT_{ex}} C_{ul} + \frac{1}{e^{h\nu/kT_{ex}} - 1} A_{ul}}{C_{ul} + \left( \frac{1}{e^{h\nu/kT_{ex}} - 1} + 1 \right) A_{ul}}$$

From detailed balance,

$$n_e (C_{eu} + B_{eu} u_\nu) = n_u (C_{ue} + B_{ue} u_\nu + A_{ue})$$

From the hint, assume collisions dominate, because those relations must always hold.

Neglect Einstein coefficients

$$n_e C_{eu} = n_u C_{ue}$$

but

$$\frac{n_u}{n_e} = \frac{g_u}{g_e} e^{-h\nu/kT_{ex}}$$

We know that in high densities when collisions dominate  $T_{ex} \rightarrow T_k$ , so

$$\textcircled{1} \frac{C_{eu}}{C_{ue}} = \frac{g_u}{g_e} e^{-h\nu/kT_k}$$

Also,

$$A_{ue} = \frac{8\pi h\nu^3}{c^3} B_{ue} \Rightarrow B_{ue} = \frac{c^3}{8\pi h\nu^3} A_{ue} \quad \textcircled{2}$$

$$\textcircled{3} g_u B_{ue} = g_e B_{eu}$$

$$u_\nu = \frac{4\pi}{c} \frac{2h\nu^3}{c^2} \frac{1}{e^{-h\nu/kT_k} - 1} = \frac{8\pi h\nu^3}{c^3} \frac{1}{e^{-h\nu/kT_k} - 1}$$

Rearranging detailed balance

$$\frac{n_u}{n_l} = \frac{C_{ul} + B_{ul} n_\nu}{C_{lu} + B_{lu} n_\nu + A_{ul}}$$

Plug in (1) + (3)

$$\frac{n_u}{n_l} = \frac{C_{ul} \frac{g_u}{g_l} e^{-h\nu/kT_u} + \frac{8\pi h\nu^3}{c^3} \frac{1}{e^{-h\nu/kT_r} - 1} \frac{g_u}{g_l} B_{ul}}{C_{ul} + \frac{8\pi h\nu^3}{c^3} \frac{1}{e^{-h\nu/kT_r} - 1} B_{ul} + A_{ul}}$$

Plug in (2)

$$\begin{aligned} \frac{n_u}{n_l} &= \frac{C_{ul} \frac{g_u}{g_l} e^{-h\nu/kT_u} + \frac{8\pi h\nu^3}{c^3} \frac{1}{e^{-h\nu/kT_r} - 1} \frac{c^3}{8\pi h\nu^3} \frac{g_u}{g_l} A_{ul}}{C_{ul} + \frac{8\pi h\nu^3}{c^3} \frac{1}{e^{-h\nu/kT_r} - 1} \frac{c^3}{8\pi h\nu^3} \frac{g_u}{g_l} A_{ul} + A_{ul}} \\ &= \frac{g_u}{g_l} \frac{C_{ul} e^{-h\nu/kT_u} + \frac{1}{e^{-h\nu/kT_r} - 1} A_{ul}}{C_{ul} + \frac{1}{e^{-h\nu/kT_r} - 1} A_{ul} + A_{ul}} \end{aligned}$$

$$\Rightarrow e^{-h\nu/kT_{ex}} = \frac{e^{-h\nu/kT_u} C_{ul} + \frac{1}{e^{-h\nu/kT_r} - 1} A_{ul}}{C_{ul} + \left( \frac{1}{e^{-h\nu/kT_r} - 1} + 1 \right) A_{ul}}$$

$$3c) I_\nu = (I_{\nu, \text{source}} + I_{\nu, \text{BG}}) e^{-\tau_\nu} + S_\nu (1 - e^{-\tau_\nu})$$

$$F_\nu = \Omega I_\nu$$

b) Case 1:

$$I_{\nu, w} = (I_{\nu, \text{source}} + I_{\nu, \text{BG}}) e^{-\tau_w} + B_\nu(T_w) (1 - e^{-\tau_w})$$

↳ for passing through warm slab only. I assumed LTE ( $S_\nu \rightarrow B_\nu(T)$ )

$$I_\nu = I_{\nu, w} e^{-\tau_c} + B_\nu(T_c) (1 - e^{-\tau_c})$$

$$= (I_{\nu, \text{source}} + I_{\nu, \text{BG}}) e^{-(\tau_w + \tau_c)} + B_\nu(T_w) (1 - e^{-\tau_w}) e^{-\tau_c} + B_\nu(T_c) (1 - e^{-\tau_c})$$

Flux would just be above expression times  $\Omega$ .

Case 2 is just Case 1

but with  $\tau_c \leftrightarrow \tau_w$  and  $T_c \leftrightarrow T_w$

$$c) I_\nu = I_{\nu, \text{BG}} e^{-(\tau_w + \tau_c)} + B_\nu(T_w) (1 - e^{-\tau_w}) e^{-\tau_c} + B_\nu(T_c) (1 - e^{-\tau_c})$$

for Case 1

Flux again  $I\Omega$

$$\Delta I = I^{on} - I^{off}$$

$$= (I_{r, source} + I_{r, DG}) e^{-(\uparrow_{\omega} + \uparrow_c)} - I_{r, DG} (e^{-(\uparrow_{\omega} + \uparrow_c)})$$

$$= I_{r, source} e^{-(\uparrow_{\omega} + \uparrow_c)}$$

$$\Rightarrow \uparrow_{\omega} + \uparrow_c = -\ln \left( \frac{\Delta I}{I_{r, source}} \right)$$