

ISM

HW #7

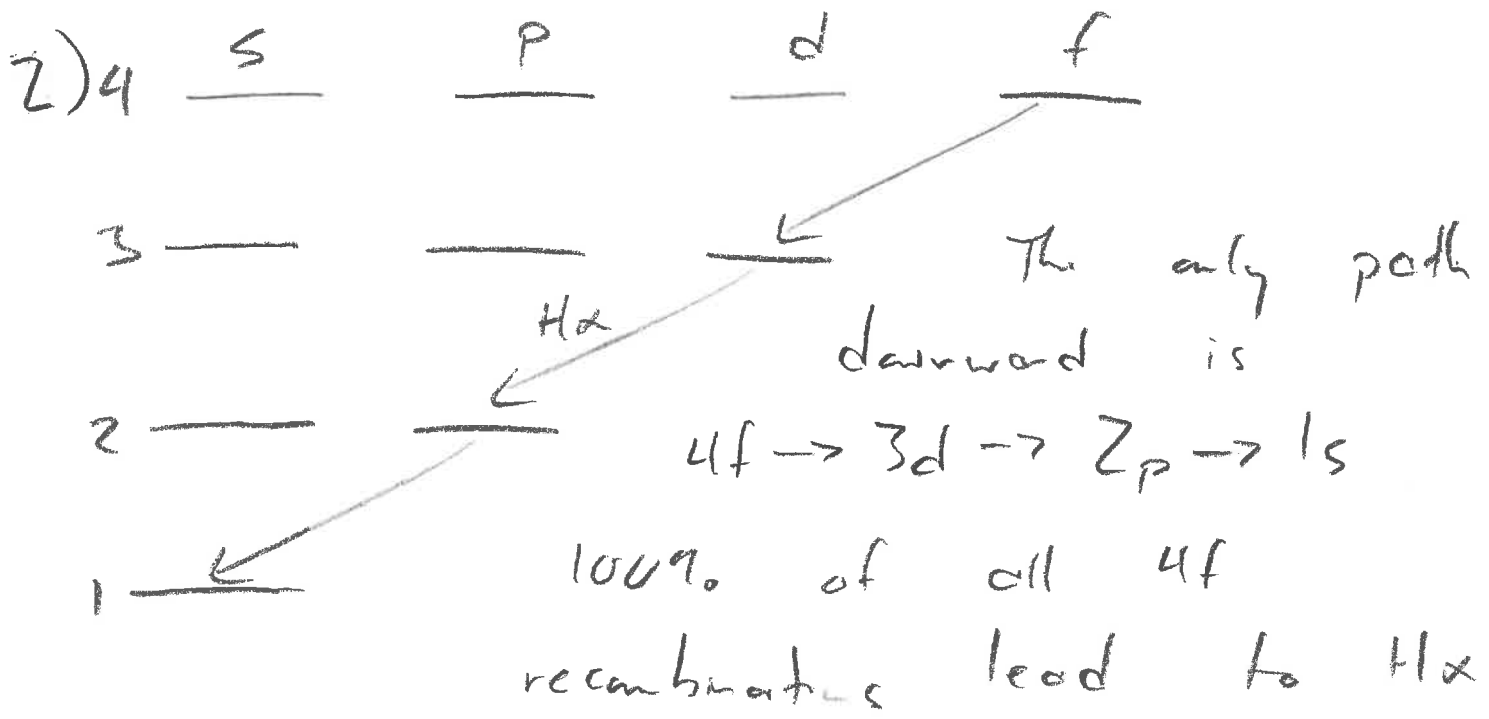
1) for such a relatively high density environment, Case B is better.
The probability is then

$n_e \alpha_B$ with $\alpha_B = 7.01 \times 10^{-12} \text{ cm}^3 \text{ s}^{-1}$
from Table 14.6

$$n_e = 0.01 \text{ cm}^{-3} \Rightarrow n_e \alpha_B = 7.01 \times 10^{-14} \text{ s}^{-1}$$

b) now we want α_{rr} , which from Table 14.6 is $\alpha_{rr} = 8.63 \times 10^{-12} \text{ cm}^3 \text{ s}^{-1}$
Recombinations are with electrons, so

$$n_e \alpha_{rr} = 0.01 \times 8.63 \times 10^{-12} = 8.63 \times 10^{-14} \text{ s}^{-1}$$



($\Delta l = 0, \pm 1$ for electric dipole)

b) We can produce $H\beta$ via $4d \rightarrow 2p$, $4p \rightarrow 2s$, or $4s \rightarrow 2p$. None of these are accessible from $4f$, so 0%.

c) We could just add them up, but it's easier to subtract from α_A , so $\alpha_{out} = \alpha_A - \alpha_{in}$. The shells fill up as

$1s \ 2s \ 2p \ 3s \ 3p \ 4s \ 3d \ 4p \ 5s \ 4d \ 5p \ 6s \ 4f$

add all these α s

We don't have α_{5s} , α_{5p} , or α_{6s} , but they are probably small.

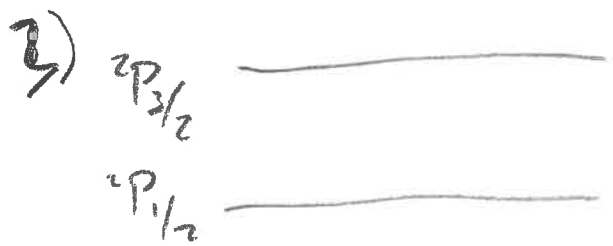
Assume $T = 10^4 K$, I get for Case A

$$\alpha_{\text{uf}} = 1.08 \times 10^{-13} \text{ cm}^3 \text{ s}^{-1}$$

($\alpha_A = 4.18 \times 10^{-13}$, so $\sim 75\%$ are to levels $< 4f$)

For case B: $\alpha_{\text{uf}} = 1.07 \times 10^{-13} \text{ cm}^3 \text{ s}^{-1}$

(these should of course be the same)



$$\lambda = 12.8 \mu\text{m} \rightarrow h\nu = 1.55 \times 10^{-22} \text{ J}$$

$$E/k =$$

From extension of our ionization discussion, the collision rate $C \hat{=} \sigma v$. At this point we could use the geometry σ and estimate the velocities from the MTS distribution.

We can do better though. Draine eqn. 2.26 says for transitions between states v and l

$$\langle \sigma v \rangle_{v \rightarrow l} = \frac{8.629 \times 10^{-8}}{\sqrt{T_4}} \frac{\Omega_{vl}}{g_v} \text{ cm}^3 \text{ s}^{-1}$$

where Ω_{vl} is the collision strength and g_v is the degeneracy.

$$\Omega_{vl} = 0.314 T_4^{0.076 + 0.002 \ln T_4} \quad (\text{assume LTE: } \Omega_{vl} = \Omega_{lv})$$

$$g_v = 2J+1 = 4$$

$$\Rightarrow \langle \sigma v \rangle_{v \rightarrow l} \hat{=} \frac{8.629 \times 10^{-8}}{\sqrt{T_4}} \frac{0.314}{4} = \frac{6.8 \times 10^{-9}}{\sqrt{T_4}} \text{ cm}^3 \text{ s}^{-1}$$

$$\langle \sigma v \rangle_{l \rightarrow v} = 1.4 \times 10^{-9} T_4^{-0.5} \text{ cm}^3 \text{ s}^{-1}$$

b) In detailed balance,

$$\frac{n_u}{n_l} = \frac{g_u}{g_l} e^{-h\nu/kT} \quad \left[\begin{array}{l} \text{assume high} \\ \text{density so ignore } A_{ul} \end{array} \right]$$

$$g_u = 4 \quad g_l = 2$$

$$\frac{n_u}{n_l} = 2 e^{-1.55 \times 10^{-22} \text{ J} / kT} \approx 2$$

back to a) for de-excitation,

$$R_{ul} = R_{lu}$$

$$g_l = 2$$

$$\langle \sigma \nu \rangle_{l \rightarrow u} = \frac{8.629 \times 10^{-5} \text{ F}}{\sqrt{T_4}} \frac{0.314}{2}$$

$$= \frac{1.3 \times 10^{-8}}{\sqrt{T_4}}$$

or
$$\frac{n_u}{n_l} = \frac{\nu_{lu}}{\nu_{ul}} \left(1 + \frac{A_{ul}}{n_l \nu_{ul}} \right)^{-1}$$

but A_{ul} is low.

4) Want $\tau_{13.6\text{eV}} = 1$

When $\tau = 1$, $d = \text{mfp}$, so $dn\sigma = 1$

From Figure 13.1, $\sigma_{\text{p1}} \approx 5 \times 10^{-18} \text{ cm}^2$
at 13.6 eV

$$dn = 2 \times 10^{17} \text{ cm}^{-2}$$

If $d = 1 \mu\text{e}$, $n = 0.07 \text{ cm}^{-3}$. If

n is greater, use case B.

If $d = 1 \text{ k}\mu\text{e}$, $n = 7 \times 10^{-5} \text{ cm}^{-3}$

