

ASTR 702

HW #5

$$1) \sigma(E) = \frac{S(E)}{E} e^{-\frac{b}{E^{1/2}}}$$

$$f(v) dv = \left(\frac{m}{2\pi kT} \right)^3 4\pi v^2 e^{-\frac{mv^2}{2kT}} dv$$

$$\langle \sigma v \rangle = \int_0^{\infty} \sigma f(v) v dv$$

$$= \int_0^{\infty} \frac{S(E)}{E} e^{-\frac{b}{E^{1/2}}} \left(\frac{m}{2\pi kT} \right)^3 4\pi v^3 e^{-\frac{mv^2}{2kT}} dv$$

$$= S(E) 4\pi \left(\frac{m}{2\pi kT} \right)^{3/2} \int_0^{\infty} \frac{v^3}{E} e^{-\left(\frac{b}{E^{1/2}} + \frac{mv^2}{2kT} \right)} dv$$

$$E = \frac{1}{2} m v^2 \quad dE = m v dv \quad dv = \frac{dE}{m v} \quad v^2 = \frac{2E}{m}$$

$$\langle \sigma v \rangle = \int_0^{\infty} \frac{2E}{m} \cdot \frac{v}{E} \cdot \frac{1}{m v} e^{-\left(\frac{b}{E^{1/2}} + \frac{mv^2}{2kT} \right)} dE$$

$$= S(E) \frac{8\pi}{m^2} \left(\frac{m}{2\pi kT} \right)^{3/2} \int_0^{\infty} e^{-\left(\frac{b}{E^{1/2}} + \frac{E}{kT} \right)} dE$$

$$= \left(\frac{8}{\pi m} \right)^{1/2} \frac{S(E)}{(kT)^{3/2}} \int_0^{\infty} e^{-\left(\frac{b}{E^{1/2}} + \frac{E}{kT} \right)} dE$$

want $\frac{d}{dE} = 0$ so $\frac{\partial}{\partial E} \left[e^{-(1/2)E^{3/2} + E/kT} \right] = 0$

Wolfram alpha gives $E = \left(\frac{6kT}{7} \right)^{2/3}$

$$2) b) q_{pp} = 1.05 \times 10^{-5} \rho X^2 T_6^4$$

$$\rho = 162 \text{ g cm}^{-3}$$

$$X = 0.34$$

$$T_6 = 15.8$$

$$\Rightarrow q_{pp} = 12.2 \text{ erg g}^{-1} \text{ s}^{-1}$$

$$q_{\omega 0} = 8.24 \times 10^{-24} \rho X \chi_{\omega 0} T_6^{12.9}$$

$$\chi_{\omega 0} = 0.013$$

$$\Rightarrow q_{\omega 0} = 4.21 \text{ erg g}^{-1} \text{ s}^{-1}$$

$$2c) \quad q_{pp} = q_{cno}$$

$$1.05 \times 10^{-5} \cancel{\times X} T_6^4 = 8.24 \times 10^{-24} \cancel{\times X} X_{cno} T_6^{19.9}$$

$$T_6^{15.9} = \frac{1.05 \times 10^{-5}}{8.24 \times 10^{-24}} \frac{X}{X_{cno}} = 3.33 \times 10^{19}$$

$$\Rightarrow T_6 = 16.9$$

$$T = 1.69 \times 10^7 \text{ K}$$

$$3) \frac{3}{2} kT = \frac{1}{2} \mu v^2$$

rms of M-B dist $v_{rms} = \left(\frac{3}{2} \frac{kT}{\mu} \right)^{1/2}$

We need $\frac{1}{2} \mu v^2 > \frac{e^2}{\epsilon_0 \cdot 4\pi r}$

$$\frac{1}{2} \mu \cdot 10^2 \cdot \frac{3kT}{2\mu} > \frac{e^2}{\epsilon_0 \cdot 4\pi r}$$

$$T > \frac{e^2 \cdot 4}{3k \epsilon_0 \cdot 4\pi \cdot 4\pi r}$$

$$> \frac{1.87 \times 10^{-7}}{5.57 \times 10^8} \cdot \frac{2.2 \times 10^{-7}}{r}$$

if $r \hat{=} 2 \times 10^{-15} \text{ m}$, $T > \frac{2.78 \times 10^7}{1.1 \times 10^8} \text{ K}$

$$3b) \frac{n_v}{n_{rms}} = \frac{\exp(-mv_{rms}^2 / 2kT) (10 v_{rms})^2}{\exp(-mv_{rms}^2 / 2kT) (v_{rms})^2}$$

$$= 5.7 \times 10^{-31}$$

3c) # H atoms

$$\frac{M_{\odot}}{m_p} = 1.2 \times 10^{57}$$

If temperature everywhere is $1.1 \times 10^8 \text{ K}$

$$N_{\text{react}} \hat{=} \frac{M_{\odot}}{m_p} \left(\frac{n_V}{n_{\text{rms}}} \right) \hat{=} 7 \times 10^{26}$$

