

AST 2105

HW #4

1) b) Maximum FLWMM

$$FLWMM = 21.47 T_4^{1/2} \text{ km/s}$$

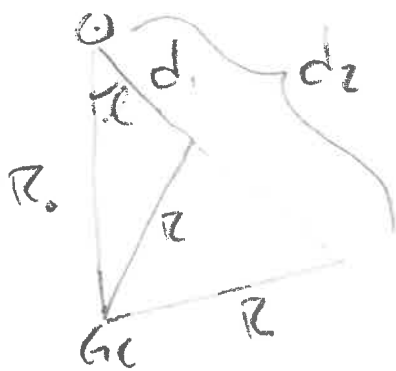
$$\Rightarrow T_4 = \left(\frac{FLWMM}{21.47 \text{ km/s}} \right)^2 \quad \text{this is } T_{\text{max}}!$$

$$V_{\text{lsr}} = R_0 \sin l \left(\frac{\Theta(R)}{R} - \frac{\Theta_0}{R_0} \right)$$

Assume $\Theta_0 = 235 \text{ km/s}$

$R_0 = 8.35 \text{ kpc}$

$$R = R_0 \left(\frac{V_{\text{lsr}}}{\Theta_0 \sin l} + 1 \right)^{-1}$$



$$R^2 = R_0^2 + d^2 - 2R_0 d \cos l$$

\Rightarrow gives 2 soln for d
for given R, l

$$Use \quad N(HI) = \frac{1.813 \times 10^{18}}{K \cdot km/s} \int \Delta T_B(v) dv$$

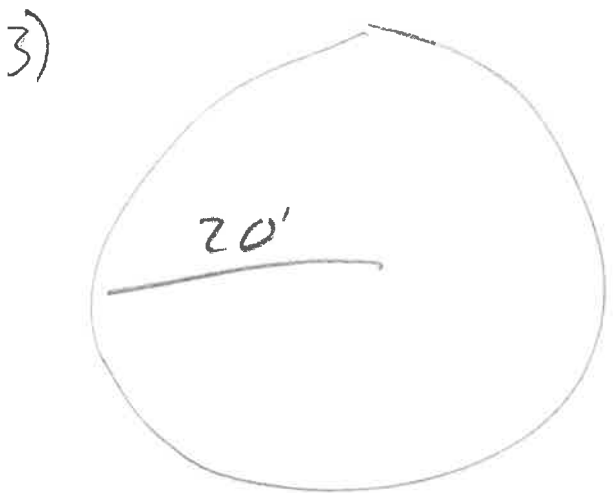
Can estimate integral from peak, FWHM of line.

$$T_S = \frac{T_A^{off} T_{RS} - T_A^{on} T_{sky}}{(T_{A_{RS}} - T_{sky}) - (T_A^{on} - T_A^{off})}$$

Look up T_{RS} in continuum data

Assume same value for T_{sky}

Since data are actually ΔT , have to work to get T_{on} , T_{off} .



Radius in pc

$$r = d\theta = 5000 \text{ pc} \cdot \frac{20' \cdot 60}{206265}$$

$$= 29 \text{ pc}$$

$$\frac{M_{\text{vir}}}{M_{\odot}} = 378 \cdot 29 \cdot 3^2 = 9.8 \times 10^4 M_{\odot}$$

$$M_{\text{LTE}} = 0.96 M_{\odot} \left[\frac{N_{T,13CO}}{8.75 \times 10^{14} \text{ cm}^{-2}} \right] \left(\frac{\theta_x}{60''} \right) \left(\frac{\theta_y}{60''} \right) \left(\frac{D}{\text{kpc}} \right)^2$$

$$N_{T,13CO} = 8.75 \times 10^{14} T_B \Delta V \text{ cm}^{-2}$$

Brightness temp is a surface brightness

so if one pointing has $T_B = 20 \text{ K}$ and all pointings are the same, the temp

above for entire cloud is $20 \text{ K} \cdot 3 \text{ kpts} = 60 \text{ K kpts}$

$$M_{\text{LTE}} = 0.96 M_{\odot} \cdot 60 \cdot 20 \cdot 20 \cdot 5^2 = 5.76 \times 10^5 M_{\odot}$$

This is telling us that the derived cloud mass is greater than the virial mass. The cloud is bound and possibly collapsing.

$$\frac{M_{\text{vir}}}{M_{\text{cl}}} = \frac{1}{5}$$

Sometimes we define the parameter

$$\alpha_{\text{vir}} = \frac{5 \sigma^2 R}{GM} = 88 \text{ for our cloud}$$