

ASTR 368

HW #5

1) a) $v = H_0 d = 7,000 \text{ km/s}$

$$H_0 = 67 \text{ km/s / Mpc}$$

$$\Rightarrow d = v / H_0 = 7,000 / 67 = 105 \text{ Mpc}$$

b) $z = \frac{v}{c} = 0.023$ non-rel.

$$z = \left(\frac{1+\beta}{1-\beta} \right)^{1/2} - 1 = \left(\frac{1+v/c}{1-v/c} \right)^{1/2} - 1 = \left(\frac{1+0.023}{1-0.023} \right)^{1/2} - 1$$
$$= 0.024 \text{ rel.}$$

c) $\frac{\Delta\lambda}{\lambda} = \frac{v}{c} \Rightarrow \Delta\lambda = 656 \text{ nm} \cdot 0.023 = 15.1 \text{ nm}$
(or 15.7 nm for rel.)

d) $m_v - M_v = 5 \log d - 5$

$$m_v = -19.3 + 5 \log (105 \times 10^6) - 5 = 15.8$$

c) $M_v = -3.53 \log P_d - 2.13 + 2.13 (B-V)$

for simplicity, assume $B-V = 0$

$$M_v = -3.53 \log 14 - 2.13 = -6.17$$

$$m_v = m_{v, \text{sr}} + 13.1 = 28.9$$

$$f) \quad d = \frac{1}{p} \quad \Rightarrow \quad p = \frac{1}{d} = \frac{1}{105 \times 10^6} = 9.5 \times 10^{-9} \text{ m}$$

Hubble resolution is $\approx 10''$

$$c) M = \frac{5 \cdot 3 \text{ Mpc} \cdot (1000 \text{ km/s})^2}{G}$$
$$= 6.9 \times 10^{45} \text{ kg} = 3.5 \times 10^{15} M_{\odot}$$

$$d) t_{\text{cross}} \sim \frac{R}{\sigma} = \frac{3 \text{ Mpc}}{1000 \text{ km/s}} = 9.26 \times 10^{16} \text{ s} = 2.93 \times 10^9 \text{ yr}$$

e) Since $t_H \sim 14 \times 10^9 \text{ yr}$, $t_H \gg t_{\text{cross}}$ and it should be virialized.

$$2) c) dU = -G \frac{M_{\text{int}} dm}{r}$$

$$U = -G \int_0^r \frac{M_{\text{int}}}{r} dm$$

$$dm = 4\pi r^2 \rho dr \quad M_{\text{int}} = \frac{4}{3}\pi r^3 \rho$$

$$\text{If } \rho \neq \rho(r), \quad U = -4\pi G \rho \cdot \frac{4}{3}\pi \rho \int_0^r r^4 dr$$

$$= -\frac{16\pi^2}{3} G \rho^2 \cdot \frac{1}{5} r^5$$

$$= -\frac{16\pi^2}{15} G \rho^2 r^5$$

$$M = \frac{4}{3}\pi R^3 \rho \Rightarrow \rho = \frac{3M}{4\pi R^3}$$

$$U = \frac{-16\pi^2}{15} G \cdot \frac{3M^2}{16\pi^2 R^6} R^5 = -\frac{3}{5} \frac{GM^2}{R}$$

$$b) 2K + U = 0$$

$$K = \frac{1}{2} M \sigma^2 \quad U = -\frac{3}{5} \frac{GM^2}{R}$$

$$M \sigma^2 = \frac{3GM^2}{5R}$$

$$M = \frac{5R\sigma^2}{3G}$$

$$\text{since } \sigma_r^2 = \frac{1}{3}\sigma^2, \quad M = \frac{5R\sigma_r^2}{G}$$