

ASTR705 ISM

HW #7

Due Wednesday, 4/5 in class

- 1) Consider an electron in an H I cloud with $n_H = 30 \text{ cm}^{-3}$, $T = 100 \text{ K}$, $n(\text{H}^+) = 0.005 \text{ cm}^{-3}$, $n(\text{C}^+) = 0.005 \text{ cm}^{-3}$, and $n_e = 0.01 \text{ cm}^{-3}$.
 - a) What is the probability per unit time for a given proton to radiatively recombine with an electron?
 - b) What is the probability per unit time for a given C^+ ion to radiatively recombine with an electron?

- 2) Suppose that an electron recombines into the $n = 4$, $\ell = 3$ ($4f$) level of hydrogen.
 - a) What is the probability that an $\text{H}\alpha$ photon will be emitted during the radiative cascade down to the ground state?
 - b) What is the probability for $\text{H}\beta$?
 - c) What is the radiative recombination rate into all levels higher in energy than $4f$ for Case A and Case B recombination?

- 3) The $12.8 \mu\text{m}$ Ne^+ line is caused by the $^2\text{P}_{1/2}$ to $^2\text{P}_{3/2}$ transition.
 - a) What is the collisional excitation (not ionization) and de-excitation rate of this transition? The units should be $\text{cm}^3 \text{ s}^{-1}$. You may have to assume a temperature.
 - b) What is the ratio of Ne^+ in the upper state compared to the lower state?

- 4) Under what conditions are Case A and Case B recombination appropriate? I'm looking for numbers here, not just general arguments!